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# Coupled－generalized nonlinear Schrödinger equations solved by adaptive step－size methods in interaction picture 

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#### Abstract

We extend two adaptive step－size methods for solving two－dimensional or multi－dimensional generalized nonlinear Schrödinger equation（GNLSE）：one is the conservation quantity error adaptive step－control method（RK4IP－CQE），and the other is the local error adaptive step－control method（RK4IP－LEM）．The methods are developed in the vector form of fourth－ order Runge－Kutta iterative scheme in the interaction picture by converting a vector equation in frequency domain．By simulating the supercontinuum generated from the high birefringence photonic crystal fiber，the calculation accuracies and the efficiencies of the two adaptive step－size methods are discussed．The simulation results show that the two methods have the same global average error，while RK4IP－LEM spends more time than RK4IP－CQE．The decrease of huge calculation time is due to the differences in the convergences of the relative photon number error and the approximated local error between these two adaptive step－size algorithms．


Keywords：nonlinear optics，optical propagation in nonlinear media，coupled－generalized nonlinear Schrödinger equations（C－GNLSE），adaptive step－size methods

PACS：42．65．Tg，42．25．Bs

## 1．Introduction

Coupling an ultrashort laser pulses into an optical fiber，a wealth of nonlinear effects will take place due to the disper－ sion and the nonlinear effect of the fiber．The phenomenon has been widely used in optical fiber communication，ultra－ fast optical，supercontinuum generation，optical coherence to－ mography，etc．${ }^{[1-4]}$ The propagation of the low power ultra－ short pulses in the fiber could be described by a mathematical model of the nonlinear Schrödinger equation（NLSE）which contains the group velocity dispersion（GVD）and self－phase modulation（SPM）terms．${ }^{[2]}$ To evaluate the high peak power femtosecond pulses，the generalized nonlinear Schrödinger equation（GNLSE）with the high order dispersion and non－ linear terms was adopted．${ }^{[5]}$ Although the NLSE could be analytically solved by the inverse scattering and self－similar method，${ }^{[6,7]}$ the GNLSE can only be calculated numerically．A common numerical solution to the GNLSE has been obtained by the split－step Fourier method（SSFM）．${ }^{[2]}$ In the SSFM scheme，the dispersion and nonlinearity were integrated re－ spectively in each step and the accuracy of the global error was of the second－order．${ }^{[8]}$ While the accuracy could be improved by the symmetric split－step Fourier method（S－SSFM）${ }^{[9-13]}$ or high order split－step schemes such as the fourth－order scheme of Blow and Wood，${ }^{[14]}$ the global accuracy was not better than that of the nonlinear step integration．In order to further im－

[^0]prove the accuracy，a number of methods have been developed： A Runge－Kutta in the interaction picture（RK4IP）method was extended to solve the GNLSE and stimulate the pulse prop－ agation and supercontinuum generation in the optic fiber by Hult．${ }^{[15]}$ The RK4IP exhibits a fifth order local accuracy and high calculation efficiency in the fixed step methods．The con－ servation quantity error adaptive step－control algorithm based on RK4IP（RK4IP－CQE）was introduced to solve the GNLSE by Heidt in 2009．${ }^{[16]}$ The RK4IP－CQE，an effective and accu－ rate numerical method of GNLSE solution，reduced more than $50 \%$ computational time than the local error adaptive step－ control method．${ }^{[17]}$ Furtherly，the RK4IP－CQE in frequency domain was also introduced to solve the GNLSE by Rieznik et al．${ }^{[18]}$ and the numerical results seems impressive．

Recently multimode optical fibers and micron waveg－ uides have reemerged as a viable platform to observe the novel linear and nonlinear physical phenomena．${ }^{[19-23]}$ The additional spatial degrees of freedom of the new fibers and waveguides offer further opportunities to investigate the in－ teresting phenomena and processes such as the vector soliton fission，the vector modulation instability，the intermodal mod－ ulational instability，and the soliton capture appeared．${ }^{[24-33]}$ Understanding the ultrashort pulse propagation dynamics mechanism behind these complex physics phenomena and processes needs to solve the numerical modeling of the
so-called coupled-generalized nonlinear Schrodinger equations (C-GNLSE) or the multimode generalized nonlinear Schrödinger equations (MM-GNLSE).

Although the MM-GNLSE can be easily handled with the SSFM and S-SSFM, the accuracies and $y$ efficiencies of the two methods need to be enhanced. To obtain accurate and effective numerical simulations, the fixed step method of RK4IP is extend to solve MM-GNLSE. ${ }^{[34-36]}$ In order to improve the calculation speed of FFT, the GPU-accelerated method has been adopted. ${ }^{[37]}$ Practically, an algorithm with various adaptive step sizes such as RK4IP-CQE or local error adaptive stepcontrol method used in GNLSE simulation may be more effective and useful than the fixed step method. However, to our knowledge, the adaptive step-size algorithm has not been extended to the solution of MM-GNSLE.

In this work we extend the adaptive step-size algorithm to C-GNLSE to verify the effectiveness. By mapping the CGNLSE in the normal picture into the interaction picture in frequency domain and converting the coupled equations into a vector equation, the RK4IP-CQE and the local error adaptive step-control method are introduced to solve C-GNLSE. The two adaptive step-control algorithms are used to solve CGNLSE to simulate the SC generation in high birefringence PCF. The simulation results are the same as those calculated by fixed step algorithm RK4IP, which proves the accuracies of the two algorithms. The calculation efficiency of RK4IPLEM and RK4IP-CQE are displayed and the large difference in computational time between the two algorithms at the same global error is explained.

## 2. Coupled generalized nonlinear Schrödinger equations

To simplify our numerical simulations, RK4IP-CQE and RK4IP-LEM are used for solving the two-dimensional CGNLSE in instead of MM-GNLSE. The process and the methods can be fully extended to solve the $n$-dimensional MMGNLSE with $n$ dimensions. The typical coupled generalized nonlinear Schrödinger equations (C-GNLSE) can be expressed as ${ }^{[38-40]}$

$$
\begin{align*}
& \frac{\partial A_{n}}{\partial z}+(-1)^{n+1} \delta \beta_{1} \frac{\partial A_{n}}{\partial T}+\sum_{m \geq 2} \frac{\mathrm{i}^{m-1} \beta_{m n}}{m!} \frac{\partial^{m} A_{n}}{\partial T^{m}} \\
= & \mathrm{i} \gamma_{n}\left(1+\mathrm{i} \tau_{\text {SHOCK }, n} \frac{\partial}{\partial T}\right)\left\{G\left(T, A_{n}, A_{3-n}\right)\right\}, \tag{1}
\end{align*}
$$

with

$$
\begin{aligned}
& G\left(T, A_{n}, A_{3-n}\right) \\
= & {\left[A_{n} \cdot \int_{-\infty}^{T} R_{1}(T-\tau)\left|A_{n}(\tau)\right|^{2} \mathrm{~d} \tau\right.} \\
& +A_{n} \cdot \int_{-\infty}^{T} R_{2}(T-\tau)\left|A_{3-n}(\tau)\right|^{2} \mathrm{~d} \tau
\end{aligned}
$$

$$
\begin{align*}
& +A_{3-n} \cdot \int_{-\infty}^{T} R_{3}(T-\tau)\left(A_{n} A_{3-n}^{*}+A_{3-n} A_{n}^{*}\right. \\
& \left.\left.\times \exp \left(-2 \mathrm{i}(-1)^{n+1} \Delta \beta z\right)\right) \mathrm{d} \tau\right], \tag{2}
\end{align*}
$$

where $A_{n}$ is the pulse envelope for the polarization $n$, the asterisk denotes the complex conjugate, the time $T=t-$ $z\left(\beta_{11}+\beta_{12}\right) / 2$ is in a reference frame moving at a group velocity, $\beta_{m n}=\partial \beta_{n} /\left.\partial \omega\right|_{\omega=\omega_{n}}$ is the $m$-th term of the Taylor series expansion for the propagation constant $\beta_{n}(\omega), \delta \beta_{1}=$ $\left(\beta_{11}-\beta_{12}\right) / 2, \Delta \beta=\beta_{01}-\beta_{02}, \gamma_{n}=n_{2} \omega_{0} / c A_{\text {eff }, n}, c$ is the speed of light in vacuum, $A_{\text {eff }, n}$ is the effective area for the polarization $n$, and $\omega_{0}$ is the carrier frequency. The express of shock term is $\tau_{\mathrm{SHOCK}, n}=1 / \omega_{0} . R_{1}(T), R_{2}(T)$, and $R_{3}(T)$ are the response functions of the fiber which is expressed as

$$
\begin{align*}
& R_{1}(T)=\left(1-f_{\mathrm{R}}\right) \delta(T)+f_{\mathrm{R}} f_{1}(T) \\
& R_{2}(T)=\frac{1}{3}\left(1-f_{\mathrm{R}}\right) \delta(T)+f_{\mathrm{R}} f_{2}(T), \\
& R_{3}(T)=\frac{1}{3}\left(1-f_{\mathrm{R}}\right) \delta(T)+f_{\mathrm{R}} f_{3}(T), \tag{3}
\end{align*}
$$

where $f_{\mathrm{R}}=0.18$ is the Raman response contribution to the Kerr effect. $\quad f_{1}(t)$ and $f_{3}(t)$ are related to the parallel and orthogonal Raman gain respectively, and can be measured experimentally. ${ }^{[41,42]}$ In the simulations, the expressions of $f_{1}(t), f_{2}(t)$, and $f_{3}(t)$ are $f_{1}=\left(\tau_{1}^{2}+\tau_{2}^{2}\right) / \tau_{1} \tau_{2}^{2} \times$ $\exp \left(-t / \tau_{2}\right) \sin \left(t / \tau_{1}\right), \quad f_{2}=f_{1}-2 f_{3}$, and $f_{3}=\left[\left(2 \tau_{\mathrm{b}}-\right.\right.$ $\left.\left.t) / \tau_{\mathrm{b}}^{2}\right)\right] \exp \left(-t / \tau_{\mathrm{b}}\right)$ respectively and the values of $\tau_{1}, \tau_{2}$, and $\tau_{\mathrm{b}}$ are $12.2 \mathrm{fs}, 32 \mathrm{fs}$, and 96 fs respectively. The C-GNLSE shown in Eq. (1) can be converted to the form indicated in Refs. [31,32,43] if the vertical and the parallel Raman gain functions are omitted.

Equation (1) can be translated to the frequency domain by Fourier transform as follows:

$$
\begin{align*}
& \frac{\partial \tilde{A}_{n}(\omega, z)}{\partial z} \\
= & -\mathrm{i}\left[(-1)^{n+1} \delta \beta_{1}\left(\omega-\omega_{0}\right)-\sum_{m \geq 2} \beta_{m n} \frac{\left(\omega-\omega_{0}\right)^{m}}{m!}\right] \tilde{A}_{n} \\
& +\mathrm{i} \gamma_{n}\left[1+\tau_{\mathrm{SHOCK}, n}\left(\omega-\omega_{0}\right)\right] \\
& \times F T\left\{G\left(T, A_{n}, A_{3-n}\right)\right\}, \tag{4}
\end{align*}
$$

where $F T$ stands for Fourier transform, and for numerical integration, it is useful to re-express Eq. (4) into the following form

$$
\begin{equation*}
\frac{\partial \tilde{A}_{n}}{\partial z}=\tilde{\hat{D}}_{n} \tilde{A}_{n}+\tilde{\hat{N}}_{n} \tilde{A}_{n}, \tag{5}
\end{equation*}
$$

where $\tilde{\hat{D}}_{n}$ and $\tilde{\hat{N}}_{n}$ are dispersion and nonlinear operators in the frequency domain

$$
\begin{aligned}
\tilde{\hat{D}}_{n} \tilde{A}_{n}= & -\mathrm{i}\left[(-1)^{n+1} \delta \beta_{1}\left(\omega-\omega_{0}\right)\right. \\
& \left.-\sum_{m \geq 2} \beta_{m n} \frac{\left(\omega-\omega_{0}\right)^{m}}{m!}\right] \tilde{A}_{n}
\end{aligned}
$$

$$
\begin{align*}
\tilde{\hat{N}}_{n} \tilde{A}_{n}= & \mathrm{i} \gamma_{n}\left[1+\tau_{\mathrm{SHOCK}, n}\left(\omega-\omega_{0}\right)\right] \\
& \times F T\left\{G\left(T, A_{n}, A_{3-n},(-1)^{n+1} \Delta \beta\right)\right\} \tag{6}
\end{align*}
$$

## 3. Algorithm

### 3.1. Fourth-order Runge-Kutta in interaction picture method

Equation (5) can be changed into the vector form,

$$
\begin{equation*}
\frac{\partial \tilde{A}}{\partial z}=\tilde{\hat{D}} \tilde{A}+\tilde{\hat{N}} \tilde{A} \tag{7}
\end{equation*}
$$

where $\tilde{\hat{D}} \tilde{A}=\left(\tilde{\hat{D}}_{1} \tilde{A}_{1} \tilde{\hat{D}}_{2} \tilde{A}_{2}\right)$, and $\tilde{\hat{N}} \tilde{A}=\left(\tilde{\hat{N}}_{1} \tilde{A}_{1} \tilde{\hat{N}}_{2} \tilde{A}_{2}\right)$. The vector equation (7) is the normal picture form of C-GNLSE in the frequency domain. Let $\tilde{A}^{\prime}=\exp (-\tilde{\hat{D}} z) \tilde{A}$, we obtain

$$
\begin{equation*}
\frac{\partial \tilde{A}^{\prime}}{\partial z}=\exp (-\tilde{\hat{D}} z) \tilde{\hat{N}}\left(\exp \left(\tilde{\hat{D}}_{z}\right) \tilde{A}^{\prime}\right) \tag{8}
\end{equation*}
$$

The vector equation (8) has the same representation as the scalar GNLSE in Ref. [16] and can be solved by RK4IP. The specific iterative scheme of RK4IP is shown in the following vector equation:

$$
\begin{align*}
& \tilde{A}_{I}(z, \omega)=\exp \left(\frac{h}{2} \tilde{\tilde{D}}\right) \tilde{A}(z, \omega), \\
& k_{1}= \\
& \left.\exp _{2}=h \frac{h}{2} \tilde{\hat{D}}\right)\left[h \tilde{\hat{N}}\left(\tilde{A}\left(\tilde{A}_{I}+k_{1} / 2\right)\left[\tilde{A}_{I}+k_{1} / 2\right]\right)\right] \tilde{A}(z, \omega), \\
& k_{3}=h \tilde{\hat{N}}\left(\tilde{A}_{I}+k_{2} / 2\right)\left[\tilde{A}_{I}+k_{2} / 2\right] \\
& k_{4}=h \tilde{\hat{N}}\left(\exp \left(\frac{h}{2} \tilde{\tilde{D}}\right)\left(\tilde{A}_{I}+k_{3}\right)\right) \\
& \quad \times \exp \left(\frac{h}{2} \tilde{\hat{D}}\right)\left[\tilde{A}_{I}+k_{3}\right] \\
& \tilde{A}(z+h, \omega) \\
& =\exp \left(\frac{h}{2} \tilde{\hat{D}}\right)\left[\tilde{A}_{I}+k_{1} / 6+k_{2} / 3+k_{3} / 3\right]+k_{4} / 6 . \tag{9}
\end{align*}
$$

The final solution of C-GNSLE can be obtained by the integration of each step using the iterative scheme of RK4IP at the fifth-order local accuracy. However, the error of RK4IP method in integration step cannot be predetermined unless the step is small enough. In order to control the error and change the step within the error in the stimulations, the local error adaptive step method and conservation quantity error adaptive step method based on RK4IP will be introduced.

### 3.2. RK4IP-LEM

For the C-GNSLE in Eq. (8), if the complex field is discretized into the frequency grid points and an integration method with RK4IP in Eq. (9) is uesed, there exists a constant for each grid point so that the calculated field can be expressed as

$$
\tilde{A}_{\text {calc }}(z+h, \omega)=\tilde{A}_{\text {ture }}(z+h, \omega)+k(\omega) h^{\eta}+o\left(h^{\eta+1}\right),(10)
$$

where $\tilde{A}_{\text {ture }}(z+h, \omega)$ is an exact solution. Here, $\eta=5$ for RK4IP scheme, for it exhibits a five-order local error. The relative local amplitude error is now defined as

$$
\begin{equation*}
\delta_{A}=\frac{\left\|\tilde{c}_{\text {calc }}-\tilde{A}_{\text {ture }}\right\|}{\left\|\tilde{A}_{\text {ture }}\right\|}, \tag{11}
\end{equation*}
$$

where $\|A\|=\left[\int\left(\left|A_{1}(\omega)\right|^{2}+\left|A_{2}(\omega)\right|^{2}\right) \mathrm{d} \omega\right]^{1 / 2}$ is the norm of vector $A$. Obviously, $\delta_{A}$ is dependent on a power of the step size $h^{\eta}$, and it can be restricted to a certain range ( $\delta_{\mathrm{G}}, 2 \delta_{\mathrm{G}}$ ) by adjusting the step size $h$ and the global error can also be controlled under a predefined goal error.

During the calculation, $\tilde{A}_{\text {ture }}$ is unknown, therefore, the algorithm estimates the local error by first taking a full step to compute a coarse solution $\tilde{A}_{\text {coarse }}$ and then independently taking two half steps towards a fine solution $\tilde{A}_{\text {fine }}$ and the truly relative local error $\delta_{A}$ in Eq. (11) can be approximated by

$$
\begin{equation*}
\delta=\frac{\left\|\tilde{A}_{\text {coarse }}-\tilde{A}_{\text {fine }}\right\|}{\left\|\tilde{A}_{\text {fine }}\right\|} \tag{12}
\end{equation*}
$$

### 3.3. RK4IP-CQE

The optical photon number $P$ during the propagation can be given by

$$
\begin{equation*}
P=\sum_{n=1,2} \int n_{\mathrm{eff}, n} A_{\mathrm{eff}, n} \frac{\left|A_{n}(z, \omega)\right|^{2}}{\omega_{0}+\Delta \omega} \mathrm{d} \omega \tag{13}
\end{equation*}
$$

where $\Delta \omega=\omega-\omega_{0}$, if $n_{\text {eff }}$ and $A_{\text {eff }}$ are constant in Eq. (13), the photon number $P$ change with the propagation $z$ can be expressed as

$$
\begin{align*}
\frac{\partial P}{\partial z}= & n_{\mathrm{eff}} A_{\mathrm{eff}} \frac{\partial}{\partial z} \int \frac{\left|A_{1}(z, \omega)\right|^{2}+\left|A_{2}(z, \omega)\right|^{2}}{\omega_{0}+\Delta \omega} \mathrm{d} \omega \\
= & n_{\mathrm{eff}} A_{\mathrm{eff}} \int\left\{\left[\frac{\partial A_{1}(z, \omega)}{\partial z} A_{1}(z, \omega)^{*}\right.\right. \\
& \left.\left.\left.+\frac{\partial A_{2}(z, \omega)}{\partial z} A_{2}(z, \omega)^{*}\right)\right]+ \text { c.c }\right\} \frac{\mathrm{d} \omega}{\omega_{0}+\Delta \omega} . \tag{14}
\end{align*}
$$

With the expressions of $\partial A_{1}(z, \omega) / \partial z$ and $\partial A_{2}(z, \omega) / \partial z$ in Eq. (4) and assuming $\gamma_{1}=\gamma_{2}=\gamma$, we will obtain $\partial P / \partial z=0$, which means that the photon number $P$ is conserved during the transmission.

If $n_{\text {eff }}, A_{\text {eff }, n}$, and $\gamma_{n}$ change with the frequency, the photon number $P$ will beconserved if the nonlinear coefficient is defined as $\gamma_{n}=3 \chi_{x x x x}^{(1)} \omega \omega_{0} / 8 c^{2} \beta_{n}(\omega) A_{\text {eff }, n} .{ }^{[44]}$ The detailed derivation of $\partial P / \partial z=0$ is shown in Appendix A.

Since the C-GNLSE conserves the photon number, the photon number $P(z)$ is equal to the true photon number $P_{\text {true }}(z+h)$ after one computational step

$$
\begin{equation*}
P_{\text {true }}(z+h)=P(z) \tag{15}
\end{equation*}
$$

It is therefore possible to calculate the relative photon number error as

$$
\begin{equation*}
\delta_{\mathrm{Ph}}=\frac{\Delta_{\mathrm{Ph}}}{P_{\text {ture }}} \tag{16}
\end{equation*}
$$

where

$$
\begin{align*}
\Delta_{\mathrm{Ph}}= & \left|P_{\text {calc }}(z+h)-P_{\text {ture }}(z+h)\right| \\
= & \left|P_{\text {calc }}(z+h)-P(z)\right| \\
= & \left\lvert\, \sum_{n=1,2} \int n_{\text {eff }, n} A_{\text {eff }, n} \frac{\left|A_{\text {calc }, n}(z+h, \omega)\right|^{2}}{\omega_{0}+\Delta \omega} \mathrm{d} \omega\right. \\
& \left.-\sum_{n=1,2} \int n_{\text {eff }, n} A_{\text {eff }, n} \frac{\left|A_{\text {ture }, \mathrm{n}}(z, \omega)\right|^{2}}{\omega_{0}+\Delta \omega} \mathrm{d} \omega \right\rvert\, \\
= & \mid \sum_{n=1,2} \int\left(\left.\left|n_{\text {eff }, n} A_{\text {eff }, n}\right| A_{\text {calc }, n}(z+h, \omega)\right|^{2}\right. \\
& \left.-n_{\text {eff }, n} A_{\text {eff }, n}\left|A_{\text {ture }, \mathrm{n}}(z, \omega)\right|^{2}\right) \left.\frac{\mathrm{d} \omega}{\omega_{0}+\Delta \omega} \right\rvert\, . \tag{17}
\end{align*}
$$

## 4. Results and discussion

In this section, the performance of adaptive step algorithms described in the above section is compared and discussed.

To prove the accuracy of the methods for C-GNLSE, a typical example of the supercontinuum generation in high birefringence photonic crystal fiber (PCF) is first simulated by using a constant step size of RK4IP. The input pulse is a hyperbolic secant with a full width at half maximum (FWHM) duration $T_{\text {FWHM }}=50 \mathrm{fs}$. The peak power of the input pulse is 20 kW , and the center wavelength is 680 nm . The angle between the polarization and the fast axis of the high birefringence PCF is $\pi / 4$. The length of high birefringence PCF is 0.1 m , and the Taylor expansion coefficients for the dispersion curve are taken from Martins et al. ${ }^{[39]}$ The nonlinear coefficient is $\gamma_{1}=\gamma_{2}=0.045 \mathrm{~W}^{-1} \cdot \mathrm{~m}^{-1}$. In the stimulation, the time window is 5 ps and discretized into $2^{13}$ grids.

Figure 1 illustrates a temporal and spectral evolution of the supercontinuum generation process over $0.1-\mathrm{m}$ length of high birefringence PCF with the step size of $40 \mu \mathrm{~m}$ in the stimulation using the constant step-size method of RK4IP. A logarithmic density scale is used which is truncated at -80 dB relative to the maximum value. As shown in Figs. 1(a)-1(d), the orthogonally polarized pulses traveling at different group velocities in the slow axial direction and the fast axial direction are completely separated in time after a few-mm propagation distance. After the orthogonally polarized pulses transmit apart, both of them break into a series of pulses at about 1.5 cm and 2 cm far in the slow axial direction and fast axial direction respectively, which is known as the vector soliton fission. ${ }^{[30]}$ The fundamental solitons emerge one by one in the slow axial direction and the fast axial direction and subsequently shift to longer wavelengths due to the intrapulse Raman scattering. Therefore, the energy is transferred to a narrow band resonance in the normal GVD regime, associated with the emergence of a dispersive wave. The spectra of the different red moving solitons and dispersion waves are formed as an octave supercontinuum. The simulations of the supercontinuum
generated in high birefringence PCF are similar to the results obtain by Martins et al. ${ }^{[39]}$


Fig. 1. Numerical simulation of supercontinuum generation in $0.1-\mathrm{m}-$ high birefringence PCF. (a) Temporal evolution and (b) spectral evolution of pulse along slow axis; (c) temporal evolution and (d) spectral evolution of pulse along fast axis, with retarded time frame of the reference travelling at envelope group velocity of input pulse used in panels (a) and (b).

To exhibit the errors changing with propagation distance, figure 2 shows the approximate local error and the relative photon number error between the values of the two consecutive computational steps with the pulse transmitting in the high birefringence PCF. The calculation is run with RK4IP method in the vector form in steps of $40 \mu \mathrm{~m}$. The approximated local error and the relative photon number error are calculated with Eqs. (11) and (16). As shown in Figs. 2(a) and 2(b), the approximated local error curve is similar to the relative photon number error curve, where they both have a large error peak located in a range from 0 m to 0.01 m and a small error peak located at about 0.02 m . To further understand the details of the error variance, the $y$ axes of the approximated local error curve and relative photon number error curve are enlarged as indicated in the insets, respectively, in Figs. 2(a) and 2(b). The details of two curves are also similar to each other and they both fluctuate heavily during the propagation in a distance of $0 \mathrm{~m}-0.05 \mathrm{~m}$; while the propagating distance is more than 0.05 m , the two error curves rise monotonically.

The oscillations of the two error curves ranging from 0 m to 0.05 m are mainly due to the vector soliton fission process. Many different frequency photons are created and annihilated in the vector soliton fission process, which makes the spectrum of the pulse change and the photon number error increase. At the same time, the time waveform of the pulse breaks into a
series of solitons, which increases the local error. When the propagation distance is larger than 0.05 m , the vector soliton fission and photons creation and annihilation become weak. Therefore the photon number error and local error vary slowly.


Fig. 2. Plots of propagation distance-dependent error estimations between two consecutive computational steps for (a) approximate local error (Eq. (11)), and (b) relative photon number error (Eq. (16)) in a constant step of $40 \mu \mathrm{~m}$, with inset showing enlarged section to compare the small scale errors, and error estimations depicted with the same $x$ axis to facilitate graph comparison.

From the above analysis, it can be concluded that when the waveforms in the time domain and in frequency domain change dramatically, the approximate local error curve and relative photon number error become large, so, the step should be reduced to produce a small error; otherwise, when the approximated local error and relative photon number error become small, step should be increased to reduce the simulation time.

Supercontinuum generation in a high-birefringence photonic crystal fiber (PCF) is also simulated with the same parameters by RK4IP-LEM with a relative local error $\eta_{\mathrm{G}}=10^{-6}$ and RK4IP-CQE with a relative photon number error $\eta_{\mathrm{P}}=$ $10^{-14}$. Figure 3 shows the calculation results. The output spectra calculated by three different methods are completely the same, which means that RK4IP-LEM and RK4IP-CQE are reliable.


Fig. 3. Slow and fast axis output spectra of high birefringence PCF calculation by RK4IP, RK4IP-LEM, and RK4IP-CQE methods.

To make a comparison of accuracy and efficiency between the RK4IP-LEM method and the RK4IP-CQE method,
the computational time value with the global error of the twoadaptive step-control algorithm is shown in Fig. 4. The global average error in Fig. 4 is defined as

$$
\begin{equation*}
\delta=\frac{\left\|\tilde{A}_{\mathrm{cal}}-\tilde{A}_{\mathrm{acc}}\right\|}{\left\|\tilde{A}_{\mathrm{acc}}\right\|} \tag{18}
\end{equation*}
$$

where the complex field $\tilde{A}_{\text {cal }}$ is calculated by the RK4IP-LEM method and the RK4IP-CQE method, $\tilde{A}_{\text {acc }}$ is calculated by the RK4IP in steps of $0.1 \mu \mathrm{~m}$.

It can be seen from Fig. 4 that the computational time taken by each of the RK4IP-LEM method and the RK4IPCQE method generally decreases with the global average error increasing. However, the computational time taken by the RK4IP-LEM method is much more than that by the RK4IPCQE method, which is 30 times more than that by the RK4IPCQE method when the global average error is less than $10^{-7}$, and is about 20 times more than that by the RK4IP-CQE method when the global average error is more than $10^{-7}$.


Fig. 4. Computational time taken by RK4IP-CQE and RK4IP-LEM against global average error, normalized by the time required to evaluate $10^{3} \mathrm{FFT}$, for the supercontinuum generation process in highbirefringence PCF.

The RK4IP-LEM method uses one more FFTs than the RK4IP-CQE at each step as indicated in the above section. If the step sizes of two methods in each iteration process are the same, the computational time of RK4IP-LEM will be twice longer than that of the RK4IP-CQE method. In order to explain the difference in computational time between the RK4IPLEM method and the RK4IP-CQE method, the error estimations between the two consecutive steps for the approximate local error and relative photon number error are calculated by the RK4IP method in different step sizes, furthermore, the valuations of the step size and the error between two consecutive steps in calculation process of the RK4IP-LEM method and the RK4IP-CQE method under different error limits are also calculated. The results are shown in Figs. 5-7, respectively.

Figure 5 shows the approximate local error and relative photon number error between two consecutive steps calculated by the RK4IP method in steps of $10 \mu \mathrm{~m}, 20 \mu \mathrm{~m}, 40 \mu \mathrm{~m}$, and $80 \mu \mathrm{~m}$. As shown in Fig. 5(a), the approximate local error curves are similar when the step size is $40 \mu \mathrm{~m}$ and $80 \mu \mathrm{~m}$, respectively. They both fluctuate heavily in a propagation distance range of $0 \mathrm{~m}-0.05 \mathrm{~m}$. When the pulse propagating dis-
tance is more than 0.05 m , the two error curves rise monotonically. The reason of the fluctuations can be found in Fig. 2. While the step size decreases to $10 \mu \mathrm{~m}$ and $20 \mu \mathrm{~m}$, the fluctuation between $0 \mathrm{~m}-0.05 \mathrm{~m}$ disappear, which is different from that in steps of $40 \mu \mathrm{~m}$ and $80 \mu \mathrm{~m}$. As shown in Fig. 5(b), except for different error amplitudes, the relative photon number curves are all nearly the same in steps of $80 \mu \mathrm{~m}, 40 \mu \mathrm{~m}$, $20 \mu \mathrm{~m}$, and $10 \mu \mathrm{~m}$. They all fluctuate heavily in the propagation distance of $0 \mathrm{~m}-0.05 \mathrm{~m}$ and rise monotonically for the propagation distance large than 0.05 m .


Fig. 5. Variations of (a) approximated local error (Eq. (11)) and (b) relative photon number error (Eq. (17)) with different propagation distances, calculated by RK4IP method in steps of $10 \mu \mathrm{~m}, 20 \mu \mathrm{~m}, 40 \mu \mathrm{~m}$, $80 \mu \mathrm{~m}$, with $y$ axis being logarithmic.

Figure 6 shows the variations of step size with the propagation distance in the process by using the RK4IP-LEM method and the RK4IP-CQE method under different approximate local error limits $\eta_{\mathrm{G}}$ and the relative photon number error limits $\eta_{\mathrm{P}}$ respectively. In Fig. 6(a), except small oscillations near 0.01 m and 0.02 m , the step size increases until the propagation distance reaches 0.05 m at the approximate local error limit $\eta_{\mathrm{G}}=10^{-5}$. The step size decreases monotonically after 0.05 m . The small oscillations near 0.01 m and 0.02 m disappear gradually and the step size increases monotonically before 0.05 m and then decreases monotonically as the $\eta_{\mathrm{G}}$ decreases to $10^{-6}, 10^{-7}$, and $10^{-8}$. The smaller the $\eta_{\mathrm{G}}$, the more gently the change of the step size is. It is shown in Fig. 6(b) that the step size curves with different relative photon number limits $\eta_{\mathrm{P}}=10^{-12} ; 10^{-13} ; 10^{-14}, 10^{-15}$ all oscillate heavily in a propagation distance range of $0 \mathrm{~m}-0.03 \mathrm{~m}$; the step size increases monotonically in $0.03 \mathrm{~m}-0.05 \mathrm{~m}$ and the step size decreases monotonically in $0.05 \mathrm{~m}-0.1 \mathrm{~m}$.


Fig. 6. Step sizes versus propagation distances obtained by (a) RK4IPLEM) and (b) RK4IP-CQE under different error limits.

Figure 7 is the approximated local error (Fig. 7(a)) and the relative photon number error (Fig. 7(b)) between the consecutive computational steps varying with the propagation distance in solving the C-GNLSE by the RK4IP-LEM method and the RK4IP-CQE method with different error limits. In Fig. 7(a) the approximate local errors are all under the presetting error limits at $\eta_{\mathrm{G}}$ equating $10^{-5}, 10^{-6}, 10^{-7}$, and $10^{-8}$. The fluctuations of the approximate local error curves become gentle as the change time of the step $\left(\eta_{\mathrm{G}}\right)$ decreases. In Fig. 7(b), except within the propagation distance between $0.01 \mathrm{~m}-0.02 \mathrm{~m}$, the relative photon number errors are all under the presetting error limits at $\eta_{\mathrm{P}}$ equating $10^{-12}, 10^{-13}, 10^{-14}$, and $10^{-15}$. The fluctuations of the relative photon number error are nearly the same as those as the $\eta_{\mathrm{P}}$ decreases, which means that the change of the step is independent of the value of $\eta_{\mathrm{P}}$.


Fig. 7. (a) Approximated local error of RK4IP-LEM and (b) relative photon number error of RK4IP-CQE between consecutive computational steps under different error limits.

From the above analyses, the difference between computational time taken by the RK4IP-LEM method and the computational time taken by the RK4IP-CQE method at the same global average error level in Fig. 4 can be qualitatively explained by the results of Figs. 5-7. Owing to the nonconserved qualitatively approximated local error, the approximated local error curves under different step sizes are not similar (see Fig. 5(a)), which means that convergence rate of the
approximated local error is much different from the error limit $\eta_{\mathrm{G}}$ at every step of the propagation distance. The different convergence rates make the change of step size and approximated local error in the process of RK4IP-LEM different (see Figs. 6(a) and 7(a)). The non-similarity of the change of the step size makes global average error of RK4IP-LEM not uniformly converge. However, the relative photon number error curves under different step sizes are similar because the relative photon number error is conserved quantity (see Fig. 5(b)), and the change of step size and the relative photon number in the process of RK4IP-CQE method under different error limits are similar (see Figs. 6(b) and 7(b)). The similarity of the changing of the step size makes the global average error of RK4IP-CQE method converge uniformly. The difference between the convergences of the relative photon number error and the approximated local error induces the large difference in computational time at the same global average error in Fig. 4.

## 5. Conclusions

By mapping the C-GNLSE in the normal picture into the interaction picture in the frequency domain, the conservation quantity error adaptive step-control method and the local error adaptive step-control method are developed based on a vector form of the fourth-order Runge-Kutta in interaction picture. To prove the efficiency of the adaptive step-control methods, the two adaptive step-control methods and the RK4IP method are used to simulate the SC generation in the high birefringence PCF. The calculation accuracy and efficiency for each of these two adaptive step-control methods are discussed. At the same global average error, the computational time of RK4IP-CQE has been improved 20 times compared with that of RK4IP-LEM due to the convergences of the relative photon number error and the approximated local error. The methods will be useful for simulating the vector pulses transmission and the supercontinuum generation in the nonlinear fiber and waveguides, the pulse propagating in the multimode optical fiber, and the interaction between different pulses.

## Appendix $A$

Let $A_{n}$ be the pulse envelope for the polarization $n$, then the photon number $P$ during propagation will be defined as

$$
\begin{equation*}
P=\sum_{n=1,2} \int n_{\mathrm{eff}, n} A_{\mathrm{eff}, n} \frac{\left|A_{n}(z, \omega)\right|^{2}}{\omega_{0}+\Delta \omega} \mathrm{d} \omega \tag{A1}
\end{equation*}
$$

Then the photon number change with propagation $z$ can be expressed as

$$
\begin{align*}
\frac{\partial P}{\partial z}= & \int\left\{n _ { \mathrm { eff1 } 1 } A _ { \mathrm { eff1 } } \left[\frac{\partial A_{1}(z, \omega)}{\partial z} A_{1}(z, \omega)^{*}\right.\right. \\
& \left.+\frac{\partial A_{1}(z, \omega)^{*}}{\partial z} A_{1}(z, \omega)\right] \\
& +n_{\mathrm{eff} 2} A_{\mathrm{eff} 2}\left[\frac{\partial A_{2}(z, \omega)}{\partial z} A_{2}(z, \omega)^{*}\right. \\
& \left.\left.+\frac{\partial A_{2}(z, \omega)^{*}}{\partial z} A_{2}(z, \omega)\right]\right\} \frac{\mathrm{d} \omega}{\omega_{0}+\Delta \omega} \tag{A2}
\end{align*}
$$

If nonlinear coefficients are defined as

$$
\begin{aligned}
& \gamma_{n}=3 \chi_{x x x x}^{(1)} \omega \omega_{0} / 8 c^{2} \beta_{n}(\omega) A_{\mathrm{eff}, n} \\
& \beta_{n}(\omega)=2 \pi / \lambda=n_{\mathrm{eff}, n} \omega / c
\end{aligned}
$$

and with expression of $\partial A_{1}(z, \omega) / \partial z$ and $\partial A_{2}(z, \omega) / \partial z$ in Eq. (4), the following equation is obtained:

$$
\begin{align*}
\frac{\partial P}{\partial z}= & \frac{3 \mathrm{i} \chi_{x x x x}^{(1)}}{8 c}\left\{\int \left\{F T\left\{G\left(t, A_{1}, A_{2}\right)\right\} A_{1}(z, \omega)^{*} \mathrm{~d} \omega\right.\right. \\
& +\int\left\{F T\left\{G\left(T, A_{2}, A_{1}\right)\right\} A_{2}(z, \omega)^{*} \mathrm{~d} \omega\right. \\
& -\int\left(F T\left\{G\left(t, A_{1}, A_{2}\right)\right\} A_{1}(z, \omega)^{*}\right)^{*} \mathrm{~d} \omega \\
& \left.-\int\left(F T\left\{G\left(T, A_{2}, A_{1}\right)\right\} A_{2}(z, \omega)^{*}\right)^{*} \mathrm{~d} \omega\right\} \tag{A3}
\end{align*}
$$

Calculating each integral term in Eq. (A3), the following equation are obtained:

$$
\begin{align*}
& \int F T\left\{G\left(T, A_{1}, A_{2}\right)\right\} A_{1}^{*}(\omega) \mathrm{d} \omega \\
= & \iiint_{-\infty}^{+\infty} R_{1}\left(\omega-\omega_{1}\right) A_{1}\left(z, \omega_{1}\right) A_{1}\left(z, \omega_{2}\right) A_{1}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{1}^{*}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
& +\iiint_{-\infty}^{+\infty} R_{2}\left(\omega-\omega_{1}\right) A_{1}\left(z, \omega_{1}\right) A_{2}\left(z, \omega_{2}\right) A_{2}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{1}^{*}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
& +\iiint_{-\infty}^{+\infty} R_{3}\left(\omega-\omega_{1}\right) A_{2}\left(z, \omega_{1}\right)\left\{A_{1}\left(z, \omega_{2}\right) A_{2}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right)\right. \\
& \left.+A_{1}^{*}\left(z, \omega_{2}\right) A_{2}\left(z, \omega_{2}+\omega_{1}-\omega\right) \exp (-2 \mathrm{i} \Delta \beta z)\right\} A_{1}^{*}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega  \tag{A4}\\
& \int F T\left\{G\left(T, A_{2}, A_{1}\right)\right\} A_{2}^{*}(\omega) \mathrm{d} \omega \\
= & \iiint_{-\infty}^{+\infty} R_{1}\left(\omega-\omega_{1}\right) A_{2}\left(z, \omega_{1}\right) A_{2}\left(z, \omega_{2}\right) A_{2}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{2}^{*}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
& +\iiint_{-\infty}^{+\infty} R_{2}\left(\omega-\omega_{1}\right) A_{2}\left(z, \omega_{1}\right) A_{1}\left(z, \omega_{2}\right) A_{1}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{2}^{*}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
& 024213-7
\end{align*}
$$

$$
\begin{align*}
&+\iiint_{-\infty}^{+\infty} R_{3}\left(\omega-\omega_{1}\right) A_{1}\left(z, \omega_{1}\right) A_{2}^{*}(\omega)\left\{A_{2}\left(z, \omega_{2}\right) A_{1}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right)\right. \\
&\left.+A_{2}^{*}\left(z, \omega_{2}\right) A_{1}\left(z, \omega_{2}+\omega_{1}-\omega\right) \exp (2 \mathrm{i} \Delta \beta z)\right\} \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega  \tag{A5}\\
& \int\left(F T\left\{G\left(T, A_{1}, A_{2}\right)\right\}\right)^{*} A_{1}(\omega) \mathrm{d} \omega \\
&= \iiint_{-\infty}^{+\infty} R_{1}^{*}\left(\omega-\omega_{1}\right) A_{1}^{*}\left(z, \omega_{1}\right) A_{1}^{*}\left(z, \omega_{2}\right) A_{1}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{1}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
&+\iiint_{-\infty}^{+\infty} R_{2}^{*}\left(\omega-\omega_{1}\right) A_{1}^{*}\left(z, \omega_{1}\right) A_{2}^{*}\left(z, \omega_{2}\right) A_{2}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{1}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
&+\iiint_{-\infty}^{+\infty} R_{3}^{*}\left(\omega-\omega_{1}\right) A_{2}^{*}\left(z, \omega_{1}\right)\left\{A_{1}^{*}\left(z, \omega_{2}\right) A_{2}\left(z, \omega_{2}+\omega_{1}-\omega\right)\right. \\
&\left.+A_{1}\left(z, \omega_{2}\right) A_{2}^{*}\left(z, \omega_{2}+\omega_{1}-\omega\right) \exp (2 \mathrm{i} \Delta \beta z)\right\} A_{1}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega  \tag{A6}\\
& \int\left(F T\left\{G\left(T, A_{2}, A_{1}\right)\right\} A_{2}^{*}(\omega)\right)^{*} \mathrm{~d} \omega \\
&=\iiint_{-\infty}^{+\infty} R_{1}\left(\omega-\omega_{1}\right)^{*} A_{2}^{*}\left(z, \omega_{1}\right) A_{2}^{*}\left(z, \omega_{2}\right) A_{2}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{2}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
&+\iiint_{-\infty}^{+\infty} R_{2}\left(\omega-\omega_{1}\right)^{*} A_{2}\left(z, \omega_{1}\right)^{*} A_{1}\left(z, \omega_{2}\right)^{*} A_{1}\left(z, \omega_{2}+\omega_{1}-\omega\right) A_{2}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega \\
&+\iiint_{-\infty}^{+\infty} R_{3}\left(\omega-\omega_{1}\right)^{*} A_{1}\left(z, \omega_{1}\right)^{*}\left\{A_{2}^{*}\left(z, \omega_{2}\right) A_{1}\left(z, \omega_{2}+\omega_{1}-\omega\right)\right. \\
&\left.+A_{2}\left(z, \omega_{2}\right)^{*} A_{1}\left(z, \omega_{2}+\omega_{1}-\omega\right) \exp (-2 \mathrm{i} \Delta \beta z)\right\} A_{2}(\omega) \mathrm{d} \omega_{1} \mathrm{~d} \omega_{2} \mathrm{~d} \omega . \tag{A7}
\end{align*}
$$

Let $\Omega=\omega_{1}, \Omega_{1}=\omega, \Omega_{2}=\omega_{2}+\omega_{1}-\omega$ and use the expression of $R_{n}\left(\Omega_{1}-\Omega\right)=R_{n}^{*}\left(\Omega-\Omega_{1}\right)$, then equations (A4) and (A5) will change into Eqs. (A8) and (A9), i.e.,

$$
\begin{align*}
& \int F T\left\{G\left(T, A_{1}, A_{2}\right)\right\} A_{1}^{*}(\omega) \mathrm{d} \omega \\
= & \iiint_{-\infty}^{+\infty} R_{1}^{*}\left(\Omega-\Omega_{1}\right) A_{1}(z, \Omega) A_{1}\left(z, \Omega_{2}+\Omega_{1}-\Omega\right) A_{1}^{*}\left(z, \Omega_{2}\right) A_{1}^{*}\left(z, \Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega \\
& +\iiint_{-\infty}^{+\infty} R_{2}^{*}\left(\Omega-\Omega_{1}\right) A_{1}(z, \Omega) A_{2}\left(z, \Omega_{2}+\Omega_{1}-\Omega\right) A_{2}^{*}\left(z, \Omega_{2}\right) A_{1}^{*}\left(\Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega \\
& +\iiint_{-\infty}^{+\infty} R_{3}^{*}\left(\Omega-\Omega_{1}\right) A_{2}(z, \Omega)\left\{A_{1}\left(z, \Omega_{2}+\Omega_{1}-\Omega\right) A_{2}^{*}\left(z, \Omega_{2}\right)\right. \\
& \left.+A_{1}^{*}\left(z, \Omega_{2}+\Omega_{1}-\Omega\right) A_{2}\left(z, \Omega_{2}\right) \exp (-2 \mathrm{i} \Delta \beta z)\right\} A_{1}^{*}\left(\Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega,  \tag{A8}\\
& \int F T\left\{G\left(T, A_{2}, A_{1}\right)\right\} A_{2}^{*}(\omega) \mathrm{d} \omega \\
= & \iiint_{-\infty}^{+\infty} R_{1}^{*}\left(\Omega-\Omega_{1}\right) A_{2}(z, \Omega) A_{2}\left(z, \Omega_{1}+\Omega_{2}-\Omega\right) A_{2}^{*}\left(z, \Omega_{2}\right) A_{2}^{*}\left(\Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega \\
& +\iiint_{-\infty}^{+\infty} R_{2}^{*}\left(\Omega-\Omega_{1}\right) A_{2}(z, \Omega) A_{1}\left(z, \Omega_{1}+\Omega_{2}-\Omega\right) A_{1}^{*}\left(z, \Omega_{2}\right) A_{2}^{*}\left(\Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega \\
& +\iiint_{-\infty}^{+\infty} R_{3}^{*}\left(\Omega-\Omega_{1}\right) A_{1}(z, \Omega)\left\{A_{2}\left(z, \Omega_{1}+\Omega_{2}-\Omega\right) A_{1}^{*}\left(z, \Omega_{2}\right)\right. \\
& \left.+A_{2}^{*}\left(z, \Omega_{1}+\Omega_{2}-\Omega\right) A_{1}\left(z, \Omega_{2}\right) \exp (2 \mathrm{i} \Delta \beta z)\right\} A_{2}^{*}\left(\Omega_{1}\right) \mathrm{d} \Omega_{1} \mathrm{~d} \Omega_{2} \mathrm{~d} \Omega \tag{A9}
\end{align*}
$$

Substituting Eqs. (A6)-(A9) into Eq. (A3), then the expression of $\partial P / \partial z=0$ is obtained.

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